A QBO cookbook: Sensitivity of the Quasi-Biennial Oscillation to resolution, resolved waves, and parameterized gravity waves

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¹⁴ Key Points:

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Sensitivity of the QBO to resolution, dissipation, wave forcing, and parameterized gravity waves is explored in a single framework. The QBO period can be tuned independently of its amplitude, but the vertical structure (particularly at lower levels) is harder to capture.

The influence of factors on the QBO can be related to their impact on wave-induced
 momentum fluxes in the deep tropics.

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21 Abstract

An intermediate complexity moist General Circulation Model is used to investigate the 22 sensitivity of the Quasi-Biennial Oscillation (QBO) to resolution, diffusion, tropical tro-23 pospheric waves, and parameterized gravity waves. Finer horizontal resolution is shown 24 to lead to a shorter period, while finer vertical resolution is shown to lead to a slower 25 period and to a larger amplitude in the lowermost stratosphere. More scale-selective dif-26 fusion leads to a faster and stronger QBO, while enhancing the sources of tropospheric 27 stationary wave activity leads to a weaker QBO. In terms of parameterized gravity waves, 28 broadening the spectral width of the source function leads to a longer period and a stronger 29 amplitude although the amplitude effect saturates when the half-width exceeds ~ 25 m/s. 30 A stronger gravity wave source stress leads to a faster and stronger QBO, and a higher 31 gravity wave launch level leads to a stronger QBO. All of these sensitivities are shown 32 to result from their impact on the resultant wave-driven momentum torque in the trop-33 ical stratosphere. Atmospheric models have struggled to accurately represent the QBO, 34 particularly at moderate resolutions ideal for long climate integrations. In particular, cap-35 turing the amplitude and penetration of QBO anomalies into the lower stratosphere (which 36 has been shown to be critical for the tropospheric impacts) has proven a challenge. The 37 results provide a recipe to generate and/or improve the simulation of the QBO in an at-38 mospheric model. 39

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Plain Language Summary

The most prominent mode of variability in the tropical stratosphere is the quasibiennial oscillation (QBO), characterized by easterly and westerly winds alternating sign every ~ 14 months. Only relatively recently have comprehensive models begun to simulate a QBO spontaneously, and even in these models the representation of the QBO typically suffers from biases. Here we elucidate the sensitivities of the QBO to a wide range of model parameters, and explore how these parameters affect the QBO behavior. We expect that these results will be helpful for tuning of more comprehensive models.

48 1 Introduction

The dominant mode of variability in the tropical stratosphere, the Quasi-Biennial Oscillation, consists of downward propagating easterly and westerly wind regimes, with a period typically ranging from 24 to 32 months (Baldwin et al., 2001). Although the

QBO is a tropical phenomenon, it impacts the atmospheric circulation and composition 52 globally through a variety of mechanisms. One of the earliest remote influences to be rec-53 ognized is the so-called "Holton-Tan effect" whereby the QBO modulates the strength 54 of the stratospheric polar vortex (Holton & Tan, 1980; Garfinkel et al., 2012; Anstey & 55 Shepherd, 2014; Rao et al., 2020b), and this effect is projected to intensify under climate 56 change (Rao et al., 2020c). The QBO also directly influences tropospheric variability by 57 affecting the Pacific subtropical jet (Garfinkel & Hartmann, 2011a, 2011b) and tropical 58 convection on both seasonal mean (Collimore et al., 2003; Liess & Geller, 2012; Rao et 59 al., 2020a) and subseasonal timescales (Yoo & Son, 2016; Zhang & Zhang, 2018; Mar-60 tin et al., 2019). QBO signals are also evident in temperature and in stratospheric con-61 stituents such as ozone and water vapor (Randel & Wu, 1996; Randel et al., 1998; Di-62 allo et al., 2018; Tian et al., 2019). 63

The QBO is driven by waves propagating upwards from the troposphere with pe-64 riods unrelated to (and much faster than) that of the resulting oscillation. Lindzen and 65 Holton (1968) showed how a QBO could be driven by a broad spectrum of vertically prop-66 agating waves (with phase speeds in both westward and eastward directions), in which 67 a two-way feedback between the waves and the background flow leads to oscillating winds. 68 The first part of the feedback is that the background flow modulates the propagation 69 and damping/dissipation of the waves. The second part of the feedback is that when the 70 waves experience damping or dissipation, they flux momentum to the background flow. 71 Holton and Lindzen (1972) and Plumb (1977) demonstrated that only two wave modes 72 (one with easterly and one with westerly phase speeds) are required as long as dissipa-73 tion of waves occurs near, and not solely at, the critical lines. An important implication 74 of this earlier work is that the period and amplitude of the oscillation are controlled, in 75 part, by the spectral range and amplitude of the momentum fluxed by these waves. The 76 particular waves associated with the QBO was the focus of later work, and both large-77 scale waves (especially Kelvin waves for the westerly regime) and smaller scale gravity 78 waves have been found to be crucial (Ern et al., 2014; Pahlavan et al., 2021). 79

While the first models began to successfully simulate a spontaneous QBO-like oscillation some 20 years ago (Takahashi, 1996, 1999; Scaife et al., 2000; Hamilton et al., 2001), only around five models participating in Coupled Model Intercomparison Project Phase 5 (CMIP5) spontaneously simulated it, and the majority of CMIP6 models still have no QBO (Richter et al., 2020; Rao et al., 2020a, 2020b). Even in CMIP models that

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succeed in simulating a QBO with period and amplitude relatively close to that observed, 85 the QBO winds suffer from an inability to propagate downwards to the lower stratosphere, 86 a bias also evident in models participating in the Quasi-Biennial Oscillation initiative 87 (QBOi; Bushell et al., 2020). Furthermore, the representation of the waves that funda-88 mentally drive the QBO differ dramatically among the QBO models (Holt et al., 2020), 89 with e.g., Kelvin wave activity barely evident in some models while too strong in oth-90 ers. Diversity in the representation of mixed Rossby-gravity waves, which also contributes 91 to the driving of the QBO, is even more pronounced (Holt et al., 2020). The models with 92 stronger convectively coupled waves rely less heavily on zonal mean forcing from param-93 eterized gravity waves (Holt et al., 2020). All but one of these models (the MIROC model) 94 also includes a parameterization of gravity waves (Bushell et al., 2020), as the resolved 95 waves are apparently not energetic enough to force the QBO at resolutions typically used 96 by these models. 97

The QBO is sensitive not only to the generation of resolved wave modes, but also 98 to their subsequent upwards propagation. Some of the resolved waves have a character-99 istic vertical wavelength of a few kilometers (figure 8 and 10 of Kiladis et al., 2009), and 100 hence a model with, say, a vertical resolution of a kilometer (which is typical of CMIP 101 and QBOi models in the lowermost stratosphere, Butchart et al., 2018) will not be able 102 to accurately represent its upward propagation. The net effect is that the resolved wave 103 forcing that reaches the QBO region, and hence the QBO itself, is influenced by verti-104 cal resolution (Geller et al., 2016; Anstey et al., 2016). Indeed, Holt et al. (2016) explored 105 a model with 7km horizontal resolution that included a realistic resolved wave spectrum 106 and plentiful small-scale gravity waves in the troposphere, but still required parameter-107 ized gravity waves due to a poor representation of resolved wave dissipation in the shear 108 zones, due in part to the relatively coarse vertical resolution. The fact that at least twenty 109 different CMIP and QBOi models still simulate a reasonable QBO reflects the fact that 110 these models tune the parameterized gravity waves so that the overall momentum forc-111 ing is sufficient. 112

The goal of this study is to identify and isolate the role of resolution, dissipation, resolved wave forcing, and parameterized wave forcing, for the QBO. While many of these sensitivities have been reported before, here we assess a broader range of sensitivities all within a single modeling framework. While it is possible to consider these factors in a multi-model ensemble such as QBOi or CMIP6, the wide diversity in the representation

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of these factors among the models limits the confidence with which one can ascribe changes 118 to a given cause. For example, the tropical climatology in comprehensive GCMs is (with 119 good reason) made as realistic as possible, which necessarily limits the ability to exam-120 ine how changing resolved waves impacts the QBO. It is also very difficult to perturb 121 the resolution of a comprehensive model without severely altering its climatology, given 122 the need to re-tune other scale-sensitive parameterizations. Here, we explore the role of 123 these three factors for the QBO in a single modeling framework, with the expectation 124 that results in our framework may be relevant to other models. Our hope is that these 125 results can be used to more intelligently tune other models. 126

After describing the model and the gravity wave scheme in Section 2, we document the sensitivity to resolution, the gravity wave scheme, the hyperdiffusion, and the resolved waves in Section 3. We then explain how these various perturbations to the model lead to changes in QBO periodicity and downward propagation to the lower stratosphere in Section 4. We summarize our results and conclude with an example use of the cookbook to improve the QBO of our control integration in Section 5.

¹³³ 2 A Model of an idealized Moist Atmosphere (MiMA)

We use the model of an idealized moist atmosphere (MiMA) introduced by Jucker 134 and Gerber (2017), Garfinkel et al. (2020a), and Garfinkel et al. (2020b). This model builds 135 on the aquaplanet models of Frierson et al. (2006), Frierson et al. (2007), and Merlis et 136 al. (2013). Very briefly, the model solves the moist primitive equations on the sphere, 137 employing a simplified Betts-Miller convection scheme (A. K. Betts, 1986; A. Betts & 138 Miller, 1986), idealized boundary layer scheme based on Monin-Obukhov similarity the-139 ory, a slab ocean, and the Rapid Radiative Transfer Model (RRTMG) radiation scheme 140 (Mlawer et al., 1997; Iacono et al., 2000). Please see Jucker and Gerber (2017) and Garfinkel 141 et al. (2020b) for more details. Orography, ocean zonal heat transport, and land-sea con-142 trast (i.e., difference in heat capacity, surface friction, and moisture availability between 143 oceans and continents) are specified as in Garfinkel et al. (2020b). 144

The details of the gravity wave scheme (developed by Alexander & Dunkerton, 1999) are included in the appendix. Unless otherwise indicated, all simulations in this paper were run with a triangular truncation at wavenumber 42 (T42; equivalent to a roughly 2.8° grid) with 40 vertical levels and a model top at 0.18hPa, for 38 years after discarding at least 10 years as spinup. Vertical levels in the lower stratosphere and tropical tropopause
layer are located at sigma levels 0.135, 0.112, 0.092, 0.076, 0.062, and 0.051, which leads
to a resolution of between 1.1km (if a scale height of 6km is used) and 1.3km (if a scale
height of 7km is used).

This specification allows for a reasonable mean state in the model. Figure 1a shows 153 the December though February climatology of the zonal winds in a control simulation 154 (hereafter CONTROL) at T85 resolution, and Figure 1b shows the standard deviation 155 of the winds. The model simulates a reasonable stratospheric and tropospheric mean state, 156 and robust variability in the tropical stratosphere. The mean state in the tropical strato-157 sphere suffers from a westerly bias which is even more severe at coarser resolution, how-158 ever, and this leads to the QBO in our model suffering from a too-strong westerly regime, 159 and concomitantly, too-weak an easterly regime. Gupta et al. (2020) found that such a 160 bias occurs more commonly in spectral cores, as compared to, say, finite volume. Such 161 a bias is also evident in some of the QBOi models examined by Bushell et al. (2020, see 162 their figure 2) and CMIP6 models examined by Rao et al. (2020b, see their figure 1). Fu-163 ture work should confirm whether the sensitivities found here are robust in a model which 164 does not suffer from this bias. Finally, midlatitude stationary waves, tropical precipita-165 tion, and stratospheric variability in CONTROL were found to be captured as well as 166 many CMIP models (Garfinkel et al., 2020a, 2020b; White et al., 2020; Garfinkel, White, 167 et al., 2021). As shown later, the model represents tropical wave modes realistically as 168 well. 169

We focus on the sensitivity of these key metrics of the QBO: the vertical structure 170 of its amplitude, quantified by the standard deviation of zonal mean zonal winds at 20hPa 171 and at 77hPa representing the mid- and lower-stratosphere respectively¹, and the pe-172 riodicity, quantified by the peak power of the Fourier transformed zonal mean zonal wind 173 at 27hPa. We focus on the period at 27hPa as the QBO is well-defined at this level even 174 in simulations with a weak QBO. All of these metrics are computed after first applying 175 a low-pass ninth-order Butterworth filter with a cutoff at 120 days in order to remove 176 high frequency wave-driven variability. The simulations performed, and the value of these 177 metrics for each simulation, are listed in Figure 2. Note that the correlation between the 178

¹ Such a definition can be used even in cases with a poorly defined QBO, unlike definitions which explicitly quantify wind maxima.

amplitude at 27hPa and the period across all simulations is small (0.11), while the correlation between the amplitude at 20hPa and 77hPa is 0.81. This immediately suggests greater flexibility in tuning the period independently of the overall amplitude than in tuning the vertical structure of the QBO.

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3 Survey of sensitivity to resolution, dissipation, resolved waves, and gravity waves

We first consider the sensitivity of the QBO to resolved processes, keeping the settings for the gravity wave scheme fixed, in Section 3.1. Section 3.2 then presents the sensitivity to the gravity wave scheme while keeping the numerics and boundary conditions fixed.

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3.1 Sensitivity to resolution, dissipation, and tropospheric stationary waves

Figure 3a shows the QBO in the ERA5 reanalysis (Hersbach et al., 2020; Pahla-191 van et al., 2021, the QBO is similar in other reanalyses) and Figure 3c shows the QBO 192 at T42 with 40 vertical levels in our CONTROL. At this resolution, MiMA simulates a 193 QBO similar to that observed: the period is slightly longer, but as shown later, relatively 194 small changes to the settings in the model can lead to an exact match. The standard de-195 viation of winds in the mid-stratosphere is realistic, though it is under-estimated lower 196 in the stratosphere. Too-weak QBO winds in the lower stratosphere is a common bias 197 in QBOi and CMIP6 models (Richter et al., 2020; Rao et al., 2020a; Bushell et al., 2020), 198 and the factors that lead to its amelioration will be discussed shortly. 199

If the number of vertical levels is increased by a factor of 3, with the extra levels 200 added in-between the existing levels while the model lid is kept fixed for a vertical res-201 olution of approximately 400m in the lowermost stratosphere, the QBO period length-202 ens to 4.1 years (consistent with the lengthening of the period found in the model of Anstey 203 et al., 2016), while the standard deviation in the lowermost stratosphere, but not near 204 20hPa, increases by more than $\sim 50\%$ (Figure 3e; similar to the effect in the model of 205 Geller et al., 2016). A *decrease* in the number of vertical levels has an opposite effect (Fig-206 ure 3b): a shorter period and a degradation in the standard deviation in the lowermost 207 stratosphere, though the standard deviation in the mid-stratosphere is unaffected. These 208 changes are summarized in Figure 4ab, which shows that both the standard deviation 209

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in the lowermost stratosphere and the period increase monotonically as vertical resolu-210 tion is increased. If the horizontal resolution is increased to T63 or T85 (Figure 5, roughly 211 equivalent to a grid of 1.9° or 1.4°), the period decreases to 1.75 years and 1.2 years re-212 spectively. The amplitude increases for the T63 integration (consistent with Giorgetta 213 et al. (2006)), but then decreases as the resolution is further increased to T85 (Giorgetta 214 et al., 2006, did not consider T85 and we are not aware of any other relevant study). These 215 changes are summarized in Figure 4cd: the period decreases monotonically as horizon-216 tal resolution is increased, while the amplitude changes are less clear. 217

Models also differ in how they specify horizontal diffusion (Table 7 of Butchart et 218 al., 2018), and early modeling studies found sensitivity to this parameter (Takahashi, 1996). 219 In our pseudo-spectral model, the order n of the hyperdiffusion operator $\kappa \nabla^n$ governs 220 the extent to which the diffusion is scale-selective. Larger n leads to greater scale-selectivity, 221 and a smaller impact of diffusion on the large scale features. The net effect is that wavenum-222 bers above the smallest resolved scale (i.e., 40 or 41 for T42) are damped more strongly 223 if the damping order n is, say, 6 (i.e., ∇^6 hyperdiffusion) than if n = 10. The CON-224 TROL hyperdiffusion is ∇^8 , and we explore sensitivity to n = 6 and n = 10 in Fig-225 ure 4ef; in all cases, we modify the hyperdiffusion coefficient κ such that the damping 226 of the highest resolved wavenumber (42 at T42) is fixed so as to not impact the numer-227 ical stability of the model. Lowering n to 6 or raising it to 10 has a strong impact on the 228 QBO amplitude: a lower value of n leads to a weaker QBO with an essentially unchanged 229 period (Supplemental Figure 1a and Figure 4ef), while a larger value of n leads to a stronger 230 QBO with a shorter period (Supplemental Figure 1b). This effect is due to the weaker 231 damping on small scale resolved waves for a larger value of n. 232

Next, we explore sensitivity of the QBO to tropospheric stationary waves, while 233 keeping other settings fixed. The stationary waves in CONTROL (both Kelvin and Rossby) 234 compare favorably to those observed (Garfinkel et al., 2020a, 2020b; Garfinkel, White, 235 et al., 2021), and as shown in Shamir et al. (2021) and Section 4.1, resolved tropical tran-236 sient waves are reasonable as well. In order to quantify the impact of tropospheric sta-237 tionary waves on the QBO, we remove land-sea contrast, orography, and east-west oceanic 238 heat transport (as discussed in detail in Garfinkel et al. (2020a) and Garfinkel et al. (2020b)), 239 while keeping the north-south oceanic heat transport of Jucker and Gerber (2017). The 240 resulting weakening of the stationary waves leads to a strengthening of the QBO by over 241

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 $_{242}$ 50% in both the mid-stratosphere and lower-stratosphere (zonally symmetric BC run in

Figure 2 and Supplemental Figure 1c) and also to a slight decrease in the period.

Overall, the properties of the QBO are sensitive to the treatment of resolved waves while holding the gravity wave drag fixed. Specifically, the resolution, horizontal diffusion, and stationary waves all impact the QBO.

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3.2 Sensitivity to gravity waves

We now turn our attention to the sensitivity of the QBO to the settings of the grav-248 ity wave scheme, taking CONTROL with T42 and 40 levels as the starting point. One 240 of the tunable parameters in the Alexander and Dunkerton (1999) GW scheme (and in-250 deed of any GW scheme) is the spectral width of the forced gravity waves (c_w in equa-251 tion A1). If c_w is decreased, then the gravity waves launched in the scheme will have a 252 narrower range of phase speeds. The idealized models of Holton and Lindzen (1972) and 253 Plumb (1977) predict that such a narrowing of launched phase speeds will lead to a de-254 crease in the amplitude of the QBO winds. We now test this prediction here. In CON-255 TROL, $c_w = 35$ m/s, and we explore sensitivity to changing this parameter in Figure 256 4gh and Supplemental Figure 2. Note that c_w is only changed from 10S to 10N (i.e. $c_w =$ 257 35m/s outside of the tropics) so as to not directly impact the representation of the mid-258 latitude and polar stratosphere and minimally impact polar downwelling. The QBO is 259 increasingly sensitive to c_w if c_w is less than around 25m/s. For $c_w = 5$ m/s, the QBO 260 essentially disappears, and for a $c_w = 15$ m/s the QBO standard deviation is little more 261 than half of the standard deviation in the CONTROL integration and the period decreases. 262 For c_w of 25m/s or higher, however, the change in the resulting QBO is relatively smaller, 263 and it appears there is a saturation effect in the period and to a lesser degree in the am-264 plitude in the mid-stratosphere, even as the lower stratospheric amplitude continues to 265 increase (Figure 4gh). 266

An additional parameter of the gravity wave scheme in our model is B_{eq} , the total amplitude of the launched gravity wave stress in the tropics (see equation A3); again, this is a common parameter of most GW schemes. In CONTROL, B_{eq} is set to be identical to the global value B_0 (which is 0.0043Pa), but this parameter is poorly constrained by observations and models often use higher or lower values (Figure 5 of Molod et al., 2012). Figure 4ij and Supplemental Figure 3 assess sensitivity to the value of this pa-

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rameter. Lowering B_{eq} leads to a weakening of the QBO, as might be expected, with a slight decrease in the period. Increasing B_{eq} leads to a stronger QBO and to a sharper decrease in the period. That a stronger B_{eq} leads to a shorter period is consistent with Figure 1 of Geller et al. (2016), Table 2 of Rind et al. (2014), Figure 13 of Giorgetta et al. (2006), and section 3.4 of Richter et al. (2014). We find, however, that the sensitivity of the period is non-monotonic (Figure 4ij).

A final parameter of the gravity wave scheme which is poorly constrained is the 279 vertical level at which gravity waves are launched. The launch height in our setup is the 280 sigma $(\frac{p}{p_s})$, where p_s is the surface pressure) level closest to, but smaller than, 0.315, but 281 other models launch at 100hPa or even higher up (Anstey et al., 2016). Raising the launch 282 level leads to a stronger QBO, and as an example we show in Supplemental Figure 3e 283 the QBO for a launch height of sigma=0.15 and c_w in the tropics of 25m/s (as in Sup-284 plemental Figure 2c). The QBO in Supplemental Figure 3e has a larger standard devi-285 ation than in Supplemental Figure 2c (which has a launch height at sigma=0.315) in both 286 the mid- and lower- stratosphere as fewer gravity waves are filtered out before entering 287 the stratosphere (Figure 2). 288

The sensitivities of the QBO to all of these model properties are summarized in Table 1. A wide range of "tuning knobs" are available, and while in our experiments the T42L40 QBO is closest to that observed outside of the lowermost stratosphere, this was the product of extensive tuning. A higher resolution version of the model could be tuned to also reproduce the QBO period and amplitude as well, a point we return to in the discussion.

4 Making sense of the changes in period and downward propagation to the lowermost stratosphere

Section 3 demonstrated that the QBO periodicity and downward propagation to the lower stratosphere are sensitive to a wide range of model parameters. We now seek to diagnose why. We focus on the metrics included in Figure 2, specifically the periodicity and the standard deviation at 77hPa (i.e., in the lower stratosphere). This section considers not only the simulations discussed in Section 3 listed in Figure 2, but also simulations included in Garfinkel et al. (2020a) and Garfinkel et al. (2020b). As these facets of the QBO are intimately connected to the location of (pseudo-)momentum fluxes associated with resolved and parameterized waves, we first consider the generation of re solved waves.

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4.1 Generation of resolved waves

The QBO is driven in part by transient waves well resolved at T42, and hence we 307 show in Figure 6 the resolved waves in CONTROL and in ERA-5 reanalysis for zonal 308 wind at 200hPa from 15S to 15N. MiMA captures the redness of the spectrum in both 309 time and wavenumber (Garfinkel, Shamir, et al., 2021; Shamir et al., 2021). It also ex-310 hibits enhanced power near the analytically predicted dry wave modes of Matsuno (1966), 311 as is evident for Kelvin waves in the symmetric spectrum near a phase speed of 25 m/s. 312 The spectrum is qualitatively similar in all resolutions in MiMA. There are differences 313 between the observed spectrum and the spectrum in MiMA, however, and we focus on 314 these differences in Supplemental Figure 4. At all resolutions, the power is too strong 315 except for symmetric $\omega - k$ combinations near the Madden Julian Oscillation (k <10 316 and low frequencies) which MiMA lacks. Note that Figure 6 and Supplemental Figure 317 4 show the logarithm base-10 of the power. Hence a difference of 0.5 in Supplemental 318 Figure 4 means log_{10} (MiMA) $-log_{10}$ (ERA5) = 0.5, or that MiMA has a factor of $10^{.5} \sim$ 310 3x more power. The bias in MiMA approaches a factor of three for $\omega - k$ combinations 320 that are most energetic in Figure 6, however such a bias is well within the range of bi-321 ases in the QBOi models evaluated by Holt et al. (2020). 322

The spectrum closer to the base of the QBO is of more relevance for wave driving 323 of the QBO. Hence we show the resolved wave spectrum at 77hPa in Figure 7. It is ev-324 ident that the simulations with 40 vertical levels struggle to simulate the mixed Rossby-325 gravity mode (and to a lesser degree the Kelvin mode), while the simulation with 120 326 levels does capture these waves (Figure 7f vs 7h for Kelvin, and Figure 7e vs 7g for the 327 mixed mode). Hence, while resolved waves in the troposphere are similar for different 328 vertical resolutions, resolved waves higher up differ more strongly². The implications for 329 the QBO periodicity and downward propagation will be considered in section 4.2 and 330 4.3.331

 $^{^{2}}$ All resolutions suffer from too much power at 77hPa, as at 200hPa (Supplemental Figure 5).

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4.2 Explaining the QBO period

We now attempt to quantify how resolved and parameterized waves drive the differences in the period of the QBO among these simulations. In order to do so, we first consider how the QBO is driven by these waves in CONTROL and then consider how this wave-driving differs among the other experiments.

Taking CONTROL at T85 as an example, the top row of Figure 8 shows the zonal 337 wind tendency due to parameterized gravity waves and resolved waves (i.e., the Eliassen-338 Palm flux divergence or EPFD) for a westerly QBO phase in the lower stratosphere (anal-339 ogous to Figure 8 of Manzini et al. (2006), Figure 7 of Garcia and Richter (2019), and 340 Figure 13 of Holt et al. (2020)), defined as winds at 40hPa between 10m/s and 15m/s 341 stronger than climatological. The anomalous QBO winds are shown in solid brown and 342 dashed blue. Similar to these previous modeling studies, gravity waves and EPFD from 343 resolved waves are of similar importance in the lower stratosphere. Higher up, gravity 344 waves dominate the forcing. The wave forcing is concentrated in the shear zones, and 345 hence acts to propagate the anomalous QBO winds downward. The forcing is quanti-346 tatively similar but of opposite sign for the QBO phase with easterly winds in the lower 347 stratosphere (bottom row of Figure 8). 348

The forcing of the QBO and the QBO itself in Figure 8 is concentrated in the deep 349 tropics, and we now distill the relative alignment of the QBO and its forcing by com-350 puting the deep-tropical (4S-4N) averaged wave forcing due to resolved and gravity waves 351 for this integration and QBO phase (Figure 9a). The tropical zonal winds are shown in 352 black. Both the resolved and parameterized waves are crucial in providing a westerly torque 353 in the shear zone below the maximum westerlies, and hence allow for the downward prop-354 agation of the westerlies. Furthermore, both resolved and parameterized waves provide 355 an easterly torque above the maximum westerlies. This vertically oriented dipole in mo-356 mentum forcing supports the downward propagation of the QBO winds as the flux pro-357 vided by waves is localized within the QBO shear zone. 358

Figure 9b is as in Figure 9a but for the T42L80 integration. In contrast to Figure 9a, the westerly torque is evident primarily in the lowermost stratosphere and not just in the shear zones, and the resolved wave forcing in particular peaks far from the shear zone. The net wave forcing is more effectively canceled out by the vertical advection term $(w * \frac{\partial u}{\partial z}; \text{ not shown})$ leading to slow downward propagation and a longer period. The

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key point of Figure 9 is that for simulations with relatively short QBO periods (Figure
9a), the momentum flux convergence is concentrated in the shear zones, while for simulations with longer QBO periods (Figure 9b), the flux is spread out in the vertical over
a much broader region. This effect is even more pronounced for resolved wave forcing
than parameterized GW, and the net effect is that the wave forcing is less effective at
propagating the QBO downwards due to a misalignment of the wave forcing with the
maximum in wind-shear.

In order to consider this effect for all simulations we have performed, we compute 371 the difference in total wave forcing between the westerly shear zone (63hPa to 41hPa, 372 orange line on Figure 9) and the region above the QBO maximum (34hPa to 20hPa, pur-373 ple line on Figure 9). We then compare this differential zonal momentum either side of 374 41hPa to the QBO periodicity in Figure 10, with each simulation shown with a distinct 375 marker. This figure includes not only the simulations discussed earlier in this paper, but 376 also the experiments included in Garfinkel et al. (2020a) and Garfinkel et al. (2020b). 377 These two diagnostics are significantly correlated with each other (correlation of -0.62), 378 whereby simulations with stronger westerly forcing in the westerly shear zone simulate 379 a faster downward propagation and subsequently a shorter period. Results are similar 380 if we average over a narrower or broader region on either side of the QBO wind max-381 imum (not shown). The corresponding correlation for the easterly QBO regime is also 382 statistically significant though weaker (correlation is 0.43, plot not shown). While a cor-383 relation does not imply causation and the wind profiles associated with a given QBO phase 384 are not identical across different integrations, the overall effect is that a wave momen-385 tum forcing dipole with extrema on either side of the wind maximum will encourage down-386 ward propagation and a faster period. 387

The period of the QBO decreases when all tropospheric stationary waves are removed (Supplemental Figure 1c) in part due to a weakened Brewer-Dobson Circulation (BDC) and hence weaker tropical upwelling. Indeed, the correlation between $\bar{w*}$ from 4S to 4N at 27hPa with the QBO period for the integrations shown in Figure 10 is 0.34, whereby stronger upwelling leads to a longer period³. While this relationship is statis-

³ Note that the BDC depends not only on stationary waves, but also on equatorial waves (which strengthen in these simulations) and also baroclinicly generated synoptic waves in midlatitudes (Jucker & Gerber, 2017; Grise & Thompson, 2013).

tically significant, the variance in periodicity associated with the BDC is much weaker than that associated with resolution, and hence the BDC strength is not the determining factor for QBO period across all of our simulations. Indeed, if we focus on integrations at T42L40 with the gravity wave settings of CONTROL (and include all of the simulations of Garfinkel et al. (2020a) and Garfinkel et al. (2020b)), the correlation is essentially unchanged (correlation of 0.29).

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4.3 Explaining the QBO downward propagation

We now turn our attention to understanding the diversity of downward propaga-400 tion into the lower stratosphere. Figure 4ij and Supplemental Figure 3 showed that a 401 stronger flux of gravity waves leads to a larger amplitude QBO both at 20hPa and 77hPa, 402 and we now test the hypothesis that stronger resolved wave power also leads to a larger 403 amplitude QBO. We quantify the role of resolved waves for the downward propagation 404 using the total power at 200hPa (below the base of the QBO) associated with variabil-405 ity between 10m/s and 20m/s for each simulation. We choose this range of power as we 406 expect these waves to be most crucial for downward propagation in the lower stratosphere 407 where winds are weak, though results are similar if we examine, say, 5m/s to 15m/s or 408 5m/s to 20m/s. Figure 11 compares the standard deviation of zonal winds at 77hPa to 409 this resolved wave power, with each simulation indicated with a marker. There is clearly 410 a significant relationship between the two, and the correlation is 0.54; that is, a stronger 411 wave forcing is associated with a larger amplitude QBO. The correlation for the east-412 erly phase speeds between -10m/s and -20m/s is 0.34. 413

An additional perspective on downward propagation can be obtained by consid-414 ering the EPFD in the lowermost stratosphere during the QBO regime with strong west-415 erly winds near 40hPa, as we would expect enhanced resolved wave driving in the low-416 ermost stratosphere to encourage downward propagation. Figure 12 considers this ef-417 fect, and Figure 12a shows the relationship between winds in the shear zone below the 418 QBO wind maximum and the resolved wave driving lower down, for a composite of events 419 with WQBO winds in the lower stratosphere (composite definition as in Figure 9). Specif-420 ically, the ordinate shows the resolved wave EPFD near 100hPa, while the absicca shows 421 the wind anomaly at 77hPa (in the shear zone) lagged by one month (EPFD is related 422 to the time rate of change of zonal winds). There is clearly a strong relationship, and 423 simulations with stronger resolved wave EPFD also simulate deeper propagation into the 424

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lowermost stratosphere with larger westerly wind amplitudes. Wave driving by gravity
wave is also significantly correlated with downward propagation to the lowermost stratosphere (Figure 12b), however the regression coefficient for gravity waves is a factor of 9
smaller than that for resolved waves, so resolved waves seem to have a larger influence
on the downward propagation in the lowermost stratosphere. Hence we conclude that
spread in the dissipation of resolved waves leads to the spread in the ability of the QBO
to propagate downwards.

432 **5** Discussion and Conclusions

The Quasi-Biennial Oscillation is the dominant mode of variability in the tropi-433 cal stratosphere, and while the wind anomalies are confined to the tropics, it impacts 434 the atmospheric circulation and composition globally through a variety of mechanisms. 435 Most models participating in various model intercomparison projects have failed to sim-436 ulate the QBO, and even the recent CMIP6 and QBOi models that succeed in simulat-437 ing a QBO-like oscillation suffer from a wide range of biases in the QBO behavior. The 438 goal of this work is to provide a "cookbook" as to the sensitivities of the QBO to a range 439 of processes, so as to enable modeling groups to more efficiently hone their efforts towards 440 improving properties of the QBO. 441

Table 1 and Figure 4 summarize the sensitivities of the QBO. Finer horizontal res-442 olution is shown to lead to faster QBO downward propagation. Finer vertical resolution 443 is shown to lead to a longer period (if the GW settings are unchanged) and to an increased 444 amplitude in the lowermost stratosphere. An increase in the order of numerical hyper-445 diffusion leads to a shorter period and a stronger amplitude. Enhancing tropospheric sta-446 tionary waves leads to a weaker amplitude. A wider gravity wave spectral width at the 447 source level leads to a slower and a stronger QBO, but the amplitude effect saturates. 448 A stronger gravity wave stress at the source leads to a faster and stronger QBO. Launch-449 ing the gravity wave at a higher level leads to a stronger QBO. While these sensitivities 450 appear robust in our modeling framework, we suspect that they can only provide qual-451 itative guidance for other models while the quantitative details may vary. For example, 452 the regression coefficient between changes in the gravity wave stress at the source and 453 the QBO standard deviation likely depends on the specific gravity wave parameteriza-454 tion implemented in a given model. 455

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These sensitivities are shown to result from the details of the resultant wave-driven 456 zonal-wind torque in the stratosphere. The period of the QBO is sensitive to the rela-457 tive wave-driven torque directly below versus directly above the QBO wind maximum, 458 and models that simulate a dipole in total wave-driven torque, with acceleration below 459 and deceleration above, simulate a faster period (Figure 10). The amplitude of the QBO 460 is shown to be related to the magnitude of the wave momentum flux with relevant phase 461 speeds that can reach the stratosphere. More wave momentum flux, whether gravity or 462 resolved, lead to a stronger QBO in the mid-stratosphere (Figure 11 and 12). 463

Many models suffer from a too-weak amplitude bias in the lowermost stratosphere. 464 Of the various parameters that can be tuned, the only "fix" we identified that does not 465 simultaneously increase the amplitude in the mid-stratosphere was to increase vertical resolution. This result is consistent with Giorgetta et al. (2006), Geller et al. (2016), and 467 Anstey et al. (2016, among others) who also find sensitivity of the QBO to vertical res-468 olution. There are other ways of increasing the amplitude at 77hPa and simultaneously 469 the amplitude higher up, but then a bias in the lower stratosphere is replaced with a bias 470 in the mid-stratosphere; the only way we found to independently modify the amplitude 471 in the lower stratosphere separately from the mid-stratosphere is via vertical resolution. 472 Another bias that is only "fixed" with increased resolution is the duration of the west-473 erly regime as compared to the easterly regime. In observations, the easterly regime per-474 sists for longer at and above 20hPa while the westerly regime persists for longer near 77hPa 475 (Figure 3a). This effect is represented in the 120 level run (Figure 3e), but not in any 476 of the L40 runs (Figure 3c and Supplemental Figure 1-3). The amelioration of these bi-477 ases is likely related to the ability of Kelvin waves to drive the westerly regime in the 478 lowermost stratosphere if 120 levels are used, but these Kelvin waves are poorly repre-479 sented with 40 levels (Figure 6 and 7). Specifically, a strong Kelvin wave climatology when 480 120 levels are used appears to result in a near-persistent layer of westerlies in the lower 481 stratosphere that resists the downward propagation of the next easterly phase, and short-482 ens the easterly phase when it finally does penetrate. 483

In contrast, in the mid-stratosphere the wave forcing is more dependent on the parameterized GW, and thus the mid-stratospheric properties of the QBO can be modified by tuning the GW stress. We now demonstrate explicitly how retuning the gravity wave parameterization can lead to an improved QBO, taking the T42L120 CONTROL run as an example. Recall that this integration simulates a realistic downward propa-

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gation to the lowermost stratosphere and a reasonable amplitude, but the period is too 489 long. Our goal is to return the gravity waves so as to lower the period while minimally 490 modifying the amplitude. Specifically, we set B_{eq} to 6.3mPa and c_w in the tropics to 20m/s; 491 both of these changes should lead to a reduction in the period, while their impacts on 492 the amplitude should mostly cancel out (Figure 4). The resultant QBO is shown in Fig-493 ure 3f (as compared to Figure 3e). It is clear that the QBO period is substantially im-494 proved, even as the amplitude is generally the same. This experiment demonstrates how 495 the QBO cookbook provided in this paper can be used to more efficiently tune the QBO. 496

The QBO in MiMA does not converge numerically. Namely, increasing the resolution does not lead to a QBO that is more realistic as compared to observations. However the total resolved wave flux, and more importantly the details of where this flux deposits momentum, differs depending on the resolution, and the QBO is sensitive to the total flux and not just the resolved flux. This highlights the fact that the GW parameterization in models must be scale-aware, and that the properties of GW must be carefully adjusted of for each resolution.

When run with 40 vertical levels, sigma levels in the lower stratosphere and trop-504 ical tropopause layer are at 0.135, 0.112, 0.092, 0.076, 0.062, and 0.051, which leads to 505 a resolution of between 1.1km (if a scale height of 6km is used) and 1.3km (if a scale height 506 of 7km is used). Previous studies using models with such a coarse resolution typically 507 failed to simulate a QBO (Giorgetta et al., 2006; Richter et al., 2014; Anstey et al., 2016; 508 Geller et al., 2016), though Rind et al. (2014) note that such a coarser vertical resolu-509 tion still enables the spontaneous generation of a QBO, but it fails to propagate down 510 to the lower stratosphere. We speculate that we nevertheless succeed in simulating a QBO 511 because the resolved wave power spectrum in MiMA is stronger than observed at 200hPa 512 (Supplemental Figure 4) and importantly also at 77hPa (Figure 7 and Supplemental Fig-513 ure 5), and so the resolved wave forcing of the QBO is still reasonable (as quantified in 514 section 4). 515

A notable exception to the general tendency of models with poor vertical resolution to fail to simulate a QBO-like oscillation comes from the studies of Yao and Jablonowski (2013) and Yao and Jablonowski (2015). They studied the spontaneous development of a QBO-like oscillation in a dry dynamical core with no convection or gravity wave scheme. Their model nevertheless supported a QBO-like oscillation, though the period was too

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long and the downward propagation did not extend to the lower stratosphere. They found
that a spectral dynamical core supported this QBO-like oscillation more than a finite
volume dynamical core, and indeed our configuration of MiMA uses a spectral dynamical core. Our model does not simulate a semi-annual oscillation, likely due to the model
lid near 70km and the requirement that gravity waves converge their momentum within
the model domain (as discussed in the appendix).

None of our simulations simulate disruptions as extreme as those that have occurred 527 in the past five years (e.g. near 2016 in Figure 3a), though the simulations with weak 528 QBOs occasionally skip a particular phase and instead simulate a prolonged, e.g., west-529 erly phase (see the $B_{eq} = 0.0023$ simulation near year 30 in Supplemental Figure 3). 530 Hence a disruption can arise spontaneously if there is relatively weak gravity wave flux 531 leaving the troposphere, even as no external perturbations are imposed in the troposphere. 532 While such a mechanism may not be relevant for the disruption in 2015/2016 when wave 533 activity was anomalously strong (Kang et al., 2020), a weakening of the QBO under cli-534 mate change (Kawatani & Hamilton, 2013; Rao et al., 2020c) may make it more suscep-535 tible to disruptions. 536

Overall, this study shows that a wide range of parameters affect the QBO, and hence 537 we expect that biases in e.g. QBO strength or periodicity can be "fixed" in a compre-538 hensive model by carefully adjusting these parameters in parallel. This effect is demon-539 strated in Figure 3f: Figure 3f shows a remarkably realistic QBO (certainly better than 540 that in many of the CMIP models considered by Richter et al. (2020); Rao et al. (2020a) 541 and Rao et al. (2020b)), particularly in terms of its penetration into the lower strato-542 sphere, obtained by enhancing the vertical resolution and adjusting the gravity wave pa-543 544 rameterization source spectrum.

545 546

6 Appendix: Implementation of a gravity wave scheme in a model of an idealized moist atmosphere (MiMA)

Gravity waves have important global effects on the circulation, temperature structure, and composition of the atmosphere, but occur on spatial scales that are too fine to be resolved by nearly all general circulation models (Alexander et al., 2010). Gravity waves carry momentum and energy vertically in the atmosphere, and they are an important forcing term in the stratospheric momentum budget. Models must parameterize these forcing terms using information on the larger-scale wind and stability fields. Most

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gravity wave schemes share a few common attributes: a series of waves with various pos-553 sible combinations of the ground-relative phase speed and horizontal wavenumber are 554 launched, and the dissipation of the waves as a function of height is based on the con-555 cepts of "breaking" (Lindzen, 1981) due to the presence of critical lines, and "satura-556 tion" (Fritts, 1984; Dunkerton, 1989), as density decreases and gravity wave amplitude 557 grows. We parameterize gravity waves following Alexander and Dunkerton (1999), Donner 558 et al. (2011), and Cohen et al. (2013), and while the criteria for breaking and dissipa-559 tion of waves is left unchanged, we have modified the properties of the wave source. This 560 appendix documents these changes. 561

A key parameter in any parameterization of gravity waves is the distribution of stress across phase speeds, and we thus repeat the treatment of this in the parameterization of Alexander and Dunkerton (1999) (their equation 17):

$$B_0(c) = sgn(\hat{c})B_m \exp\left[-\left(\frac{c-c_0}{c_w}\right)^2 \ln 2\right]$$
(1)

Here c is the ground-relative phase speed; c_0 is the phase speed with maximum flux mag-565 nitude B_m , and in all experiments in this paper $B_m = 0.4m^2/s^2$; c_w is the half-width 566 at half-maximum of the Gaussian (35m/s in all integrations poleward of 10S and 10N), 567 and 35m/s in the tropics as well unless specified otherwise); and \hat{c} is the intrinsic phase 568 speed at source level. The source level is set at 315hPa in the tropics (following Don-569 ner et al., 2011) unless otherwise specified. The spectral resolution for the phase speed 570 bins is 2m/s, and the tropical wave spectrum is set to be symmetric about the zonal wind 571 at the source level (c_0 is set to the zonal wind), for all integrations shown in this paper. 572

 $B_0(c)$ represents the gravity wave amplitude during an active wave event, however gravity waves are by their very nature intermittent. The parameterization of Alexander and Dunkerton (1999) handles this intermittency by a separate parameter F_{S0} which is intended to represent the long-term average of momentum flux integrated across all phase speeds. F_{S0} and $B_0(c)$ are related by an intermittency factor ϵ following equation 19 of Alexander and Dunkerton (1999) as

$$\epsilon = \frac{F_{S0}\Delta c}{\bar{\rho_o}\sum_c |B_0(c)|\Delta c} \tag{2}$$

The value of F_{S0} in many GW parameterizations is not constant in latitude (Donner et al., 2011; Molod et al., 2012; Anstey et al., 2016), and we explore the importance of

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latitudinal dependence in F_{S0} as described in Equation 3:

$$F_{S0}(\phi) = \begin{cases} Bt_0 + 0.5Bt_{SH}(1. + \tanh(\frac{\phi - \phi_{0s}}{\delta\phi_s})) &, \quad \phi \le \phi_{0s} \\ Bt_0 + \frac{Bt_{eq} - Bt_0}{\phi_{0s} - \delta\phi_s}(\phi_{0s} - \phi) &, \quad \phi_{0s} \le \phi < \delta\phi_s \\ Bt_{eq} &, \quad \delta\phi_s \le \phi \le \delta\phi_n \\ Bt_0 + \frac{Bt_{eq} - Bt_0}{\phi_{0n} - \delta\phi_n}(\phi_{0n} - \phi) &, \quad \delta\phi_n < \phi \le \phi_{0n} \\ Bt_0 + 0.5Bt_{NH}(1. + \tanh(\frac{\phi - \phi_{0n}}{\delta\phi_n})) &, \quad \phi \ge \phi_{0n} \end{cases}$$
(3)

In CONTROL, $Bt_0 = 0.0043$ Pa, and $Bt_{eq} = Bt_0 = 0.0043$, such that the same stress 582 is imposed in both the tropics and subtropics, but we explore sensitivity to Bt_{eq} . Ad-583 ditional stress is included in midlatitudes and subpolar latitudes by setting $Bt_{NH} = 0.0035$ Pa 584 and $Bt_{SH} = 0.0035$ Pa; this extra drag helps to keep the polar vortex from becoming 585 too strong. Note that we do not include any orographic gravity wave drag in our model 586 setup. Finally, $\phi_{0n} = 15$, $\phi_{0s} = -15$, $\delta \phi_n = 10$, $\delta \phi_s = -10$ specify the meridional ex-587 tent of the QBO, and are also unchanged in all of our experiments. This functional form 588 loosely follows a similar form in the GEOSCCM model and MERRA-2 reanalysis (Figure 589 5 of Molod et al., 2012) and the Canadian Middle Atmosphere Model (CMAM, Anstey 590 et al., 2016). The net effect of this change is that the intermittency factor ϵ is made a 591 function of latitude, and specifically gravity waves are more frequently present in mid-592 latitudes, and also in the tropics if Bt_{eq} is larger than Bt_0 . 593

An additional change made from the configuration in Alexander and Dunkerton 594 (1999) and Cohen et al. (2013) is that the momentum associated with gravity waves that 595 would leave the upper model domain is deposited evenly in the levels above 0.85hPa in 596 order to conserve momentum. (There are three such levels when the model is run with 597 40 total levels.) This avoids any complications noted by Shepherd and Shaw (2004) and 598 Shaw and Shepherd (2007) associated with non-conservation of momentum. Note that 599 Cohen et al. (2013) inserted this momentum evenly in the levels above 0.5hPa. No sponge 600 layer is included in the model. This requirment, coupled with the lid near 70km, likely 601 kills the semi-annual oscillation in MiMA. 602

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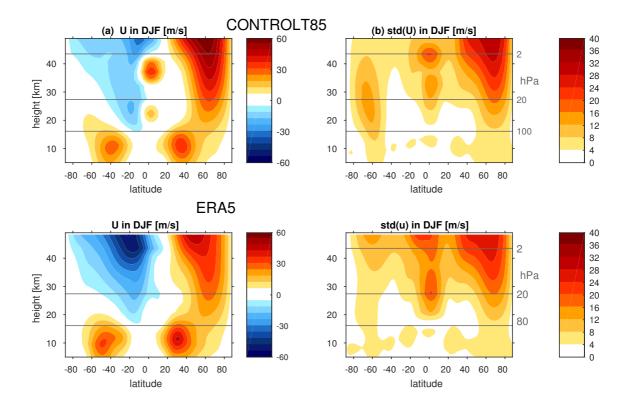


Figure 1. (a) Zonal mean zonal wind climatology in December through February; (b) standard deviation of the zonally averaged zonal wind. For (a), the contour interval is 6m/s and the 0m/s contour is omitted. (top) in Control at T85 with 40 vertical levels; (bottom) in ERA5

		ist of simulations in this paper		
	std dev 20hPa	std dev 77hPa	period 27hPa	-
ERA-5	17.9	4.5	2.4	
CONTROL	14.9	3.3	2.5	1
CONTROL 120 levels	14.5	5.5	4.1	2
CONTROL 30 level	14.4	2.7	2.0	з
CONTROL, T85	16.9	3.0	1.2	4
CONTROL, T63	18.3	3.8	1.8	5
CONTROL, nabla6	12.1	3.2	2.5	6
CONTROL, nabla10	19.3	4.2	2.2	7
zonally symmetric BC	23.9	4.9	2.3	8
CONTROL, cw=45	15.5	3.5	2.5	9
CONTROL, cw=25	13.8	2.9	2.5	10
CONTROL, cw=15	8.9	2.4	2.1	11
CONTROL, cw=5	3.7	2.1	1.0	12
CONTROL, Beq=0.0063	22.5	5.1	2.0	13
CONTROL, Beq=0.0023	10.0	2.2	2.4	14
CONTROL, Beq=0.0013	5.6	2.2	2.3	15
CONTROL, launch=150hPa, cw=25	16.5	3.4	2.4	16

Figure 2. A list of experiments included in this paper, with color shading added for clarity. Note that in addition to these 16 simulations, the scatter plots show additional integrations used in Garfinkel et al. (2020a) and Garfinkel et al. (2020b). Experiment 1 was performed at T42 with 40 vertical levels, ∇^8 hyperdiffusion, cw=35m/s, B_{eq} =0.0043Pa, and a launch height of 315hPa, and the other experiments use these settings except as otherwise specified. For ERA-5, the standard deviation at 80hPa is shown instead of 77hPa, and the period is computed at 30hPa instead of 27hPa. Note that while the T42L40 simulations simulate too weak a standard deviation at 20hPa, they simulate too strong a standard deviation at 10hPa.

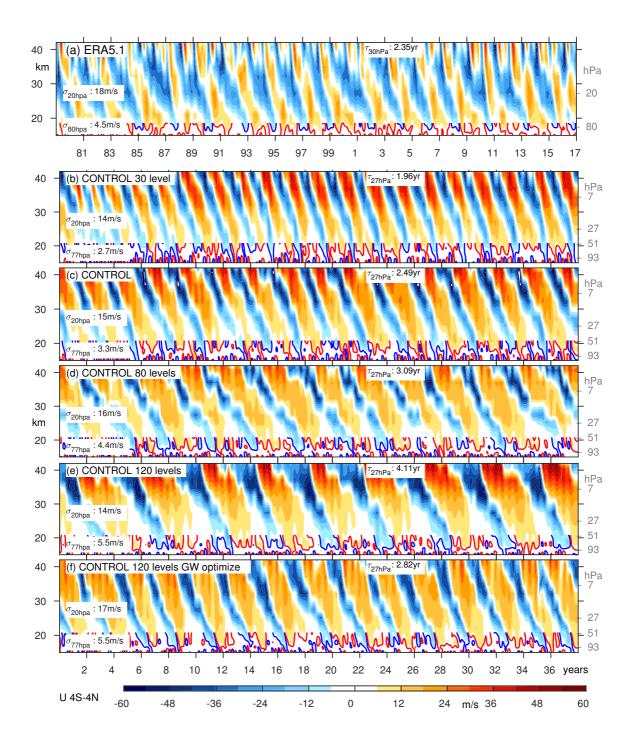


Figure 3. Zonal mean zonal wind from 4S-4N in (a) ERA5; Control at T42 with (b) 30, (c) 40, (d) 80, (e) 120 vertical levels; (f) QBO in a T42L120 run in which the gravity wave settings have been modified to improve the QBO periodicity. Specifically B_{eq} is set to 6.3mPa and c_w in the tropics to 20m/s. Each panel indicates the standard deviation of winds at 20hPa and 77hPa, and the period at 27hPa. The contour interval is 6m/s, and the 3m/s contour is shown in blue and red in the lower stratosphere.

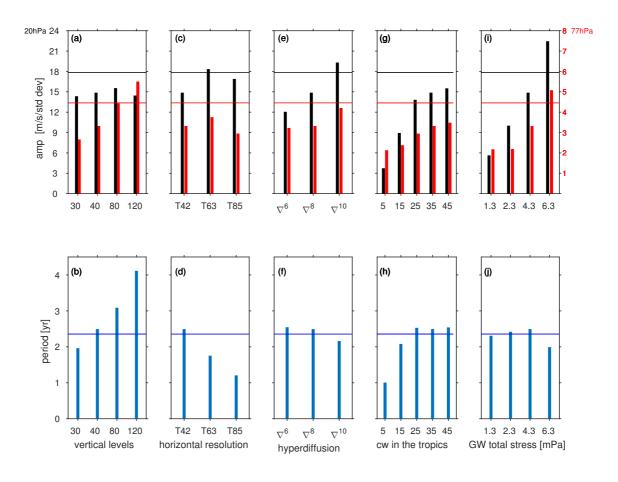


Figure 4. Summary of the sensitivities of the QBO period and amplitude to (a-b) vertical resolution; (c-d) horizontal resolution; (e-f) hyperdiffusion order; (g-h) spectral width of the launched gravity waves in the tropics; (i-j) total gravity wave stress in the tropics. A horizontal line denotes the corresponding value from ERA-5.

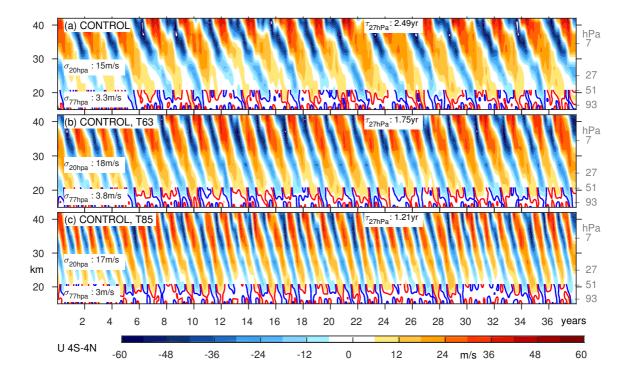


Figure 5. Zonal mean zonal wind from 4S-4N in (a) Control at T42 with 40 vertical levels; (b) Control at T63 with 40 vertical levels; (c) Control at T85 with 40 vertical levels. Each panel indicates the standard deviation of winds at 20hPa and 77hPa, and the period at 27hPa. The contour interval is 6m/s, and the 3m/s contour is shown in blue and red in the lower stratosphere.

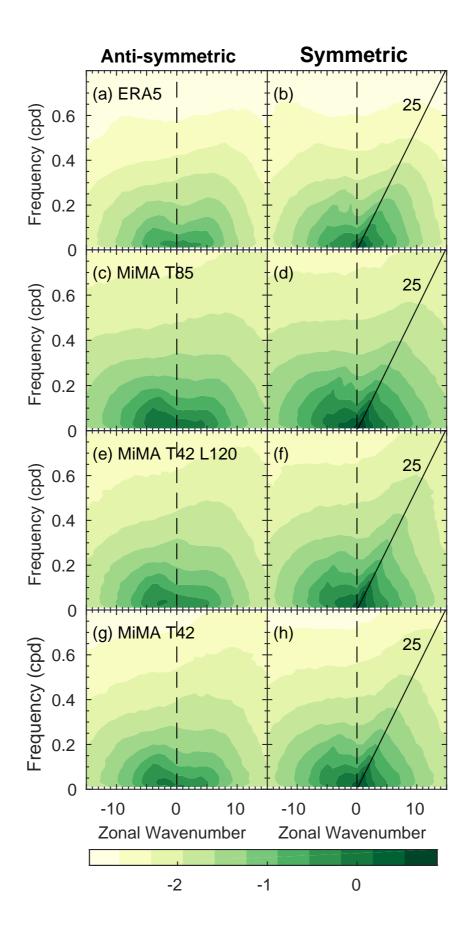


Figure 6. The logarithm base-10 of the raw 39mmetric and anti-symmetric spectrum of zonal wind at 200hPa from 15S to 15N in (a-b) ERA5; (c-d) Control at T85 with 40 vertical levels; (e-f) Control at T42 with 120 vertical levels; (g-h) Control at T42 with 40 vertical levels.

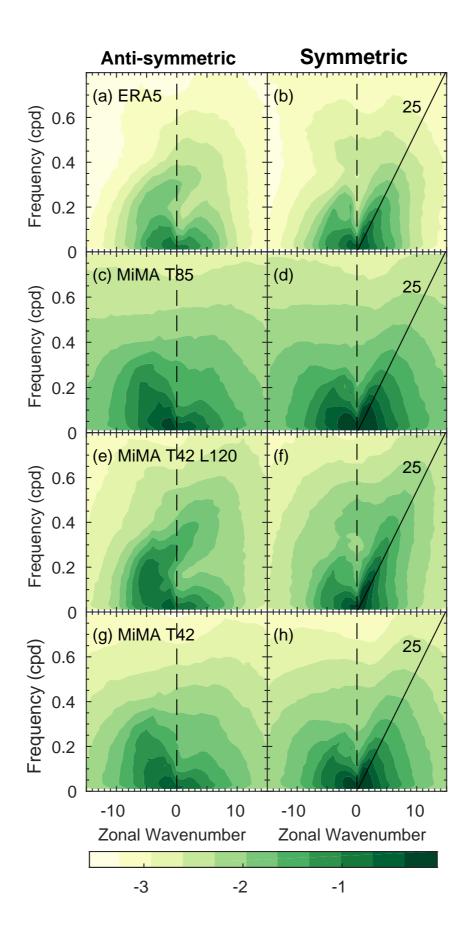


Figure 7. As in #4gure 6 but for 77hPa.

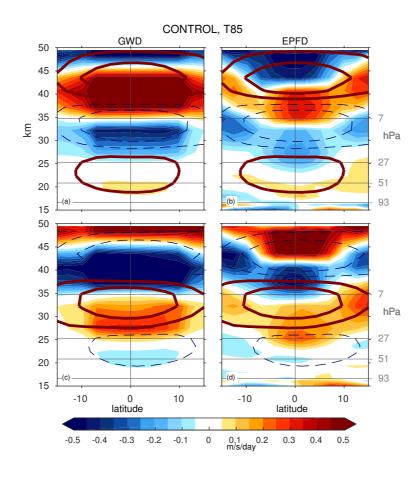


Figure 8. Forcing of winds by (left) parameterized gravity waves and (right) resolved waves for CONTROL at T85 for a QBO phase defined as wind anomalies at 41hPa between (top) 10m/s and 15m/s (i.e. WQBO) and (bottom) -10m/s and -15m/s (i.e. EQBO). Results are similar for other resolutions (not shown).

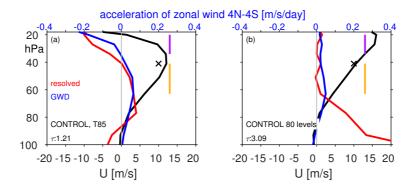


Figure 9. QBO and its resolved and parameterized wave forcing in integrations with a relatively (left) fast period and (right) slow period for a WQBO composite in which anomalous zonal winds at 41hPa must be between 10 and 15m/s. The x-axis for the QBO is shown on the bottom, and for the wave forcings on the top. Orange and purple lines show regions averaged over for Figure 10.

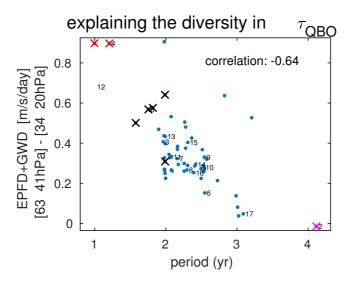


Figure 10. Relationship between QBO periodicity and the difference in total wave driving on either side of the winds at 41hPa (see orange and purple lines in Figure 9), for a WQBO composite in which anomalous zonal winds at 41hPa must be between 10 and 15m/s. Numbering of experiments follows Figure 2, and additional experiments performed as part of Garfinkel et al. (2020b) and Garfinkel et al. (2020a) are shown unnumbered for clarity. Black x-es correspond to runs at T63, red x-es to runs at T85, and magenta to runs with 120 levels.

Table 1. summary of the sensitivities of the QBO

	period	amplitude
finer horizontal resolution	faster	small effect
finer vertical resolution	slower	stronger but only in lowermost stratosphere
higher hyperdiffusion power	faster	stronger
adding tropospheric stationary waves	small effect	weaker
wider gravity wave spectral width	slower	stronger, but effect saturates
stronger gravity wave amplitude	faster	stronger
higher gravity wave launch level	small effect	stronger

Table: summary of the QBO's sensitivities

(ARC) Centre of Excellence for Climate Extremes (CE170100023). Correspondence should

be addressed to C.I.G. (email: chaim.garfinkel@mail.huji.ac.il). The updated version of

MiMA used in this study including the modified source code and example name lists to

reproduce the experiments can be downloaded from https://github.com/ianpwhite/MiMA/releases/tag/MiMA-

⁶¹² ThermalForcing-v1.0beta (with DOI: https://doi.org/10.5281/zenodo.4523199). It is ex-

pected that these modifications will also eventually be merged into the main MiMA repos-

itory which can be downloaded from https://github.com/mjucker/MiMA.

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617	flow forcing due to breaking grav	vity waves. Journal of the atmospheric
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619	ASPOMF \rangle 2.0.CO;2	

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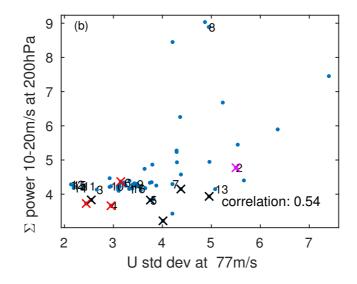


Figure 11. Relationship between QBO standard deviation at 77hPa and the resolved wave driving at 200hPa between 10m/s and 20m/s. The resolved wave driving in this range can be computed by summing over the appropriate spectral bins in, say, Figure 6. Numbering of experiments follows Figure 2, and additional experiments performed as part of Garfinkel et al. (2020b) and Garfinkel et al. (2020a) are shown unnumbered for clarity.

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635	Geophys., 39(2), 179-229.
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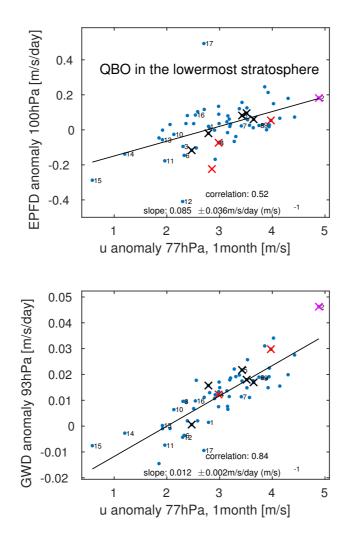


Figure 12. Relationship between winds in the shear zone below the QBO wind max and the wave driving lower down, for a WQBO composite in which anomalous zonal winds at 41hPa must be between 10 and 15m/s. Numbering of experiments follows Figure 2, and additional experiments performed as part of Garfinkel et al. (2020b) and Garfinkel et al. (2020a) are shown unnumbered for clarity.

643	Bushell, A., Anstey, J., Butchart, N., Kawatani, Y., Osprey, S., Richter, J., oth-
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